Linking ridges, arches and vertical waves in the kinematics of the Milky Way



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Substructure in the Solar neighbourhood

Local (d<200 pc) moving groups with Hipparcos.



Expect ~ Gaussian blob



Dehnen (2000)

(tangential in the direction of rotation)

[km/s]

Substructure in the Solar neighbourhood (d<200 pc)

/ [km/; (tangential Gaia DR2 (2018)



U [km/s] (radial)

Trick et al. (2018)

Diagonal Ridges in R-Vphi

- New substructure!
- Extend over several kpc



Antoja et al. (2018)

Diagonal Ridges in R-Vphi

Also prominent in maps of radial and vertical velocity



Galactocentric Radius (kpc)

Diagonal Ridges in R-Vphi

Ridge stars within 0.2 kpc from the mid-plane ?



Galactocentric Radius (kpc)

Metal rich \rightarrow young stars?



Constant energy & Lz curves



<V_R> [km/s]

<Vz> [km/s]

Galactocentric Radius



Correlations?

$(|R - R_{\odot}| < 1.0)\&(|\phi - \phi_{\odot}| < 25)$ 8 peaks in density



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Correlate in density/Z/Vz/VR





Usual suspects

Bar

Spiral arms

Trick/Monari etc

Long-lived (density wave)?

Transient structures (N-body sims)

Usual suspects

Spiral arms

Long-lived (density wave)? Transient structures (N-body sims)

J. A. S. Hunt et al.



Long slow bar + transient spirals

Usual suspects

Spiral arms

Long-lived (density wave)? Transient structures (N-body sims)



Long slow bar + transient spirals

Hunt et al. (2018)

1. Simulating R-vphi plane (phase-mixing simulation)



Initial distribution: four thin spiral arms Radial velocity N(0,20) Azimuthal N(Vc,20)

Evolve in axisymmetric potential (disc+bulge+spherical halo)

• MWPotential 2014 Galpy (Bovy 2015)

Liouville's theorem

full phase-space density (or volume) of a system evolving in a fixed potential is conserved (X,Y, VX,VY)



UV arches

Time



2. Simulating R-vphi plane (N-body simulation)





Thorsten Tepper-Garcia (USyd)

- Case 1 (isolated Galaxy); case 2 (Galaxy + interaction ex: with Sgr dwarf)
- Initial conditions: dynamical equilibrium but instabilities always present (goal)
- Positions/velocities set using DICE code, evolved in RAMSES

Isolated Galaxy



External vs. Internal ?



Resonances?

Galactic bar (length = 3-5 kpc)	Features in UV distribution
Short & Fast	Only Hercules stream
Long & slow	low velocity arches but No Hercules stream
Long & faster	Hercules, and very high velocity arches

Resonances alone (bar and/or spirals) are not enough to explain ALL the features seen by Gaia

Hunt et al. (2019)

To summarize...

Khanna et al. (2019)



- DR2 R-Vphi ridges seen in density/velocity/[Fe/H]/height.
- Simple phase-mixing models \rightarrow qualitatively reproduce.
- Transient spirals in N-body simulations \rightarrow also qualitatively reproduce
- Both resonances and phase-mixing are important.

Near Future

- Are Ages (reliable) informative?
- Test with N-body simulations initialised with distribution functions (e.g., AGAMA).
- Perhaps future Gaia releases (deeper than RVS) will reveal many more ridges?
- How are the warp, spirals and corrugations related.